

Regular Article - Experimental Physics

Disentangling the sources of ionizing radiation in superconducting qubits

L. Cardani ^{1,a}, I. Colantoni ^{1,2}, A. Cruciani ¹, F. De Dominicis ^{3,4,b}, G. D'Imperio ¹, M. Laubenstein ⁴, A. Mariani ^{1,c}, L. Pagnanini ^{3,4,5}, S. Pirro ⁴, C. Tomei ¹, N. Casali ¹, F. Ferroni ^{1,3}, D. Frolov ⁶, L. Gironi ^{7,8}, A. Grassellino ⁶, M. Junker ⁴, C. Kopas ⁹, E. Lachman ⁹, C. R. H. McRae ^{10,11,12}, J. Mutus ⁹, M. Nastasi ^{7,8}, D. P. Pappas ⁹, R. Pilipenko ⁶, M. Sisti ⁸, V. Pettinacci ¹, A. Romanenko ⁶, D. Van Zanten ⁶, M. Vignati ^{1,13}, J. D. Withrow ¹⁴, N. Z. Zhelev ¹⁵

- ¹ INFN, Sezione di Roma, P.le Aldo Moro 2, 00185 Rome, Italy
- ² Consiglio Nazionale delle Ricerche, Istituto di Nanotecnologia, c/o Dip. Fisica, Sapienza Università di Roma, 00185 Rome, Italy
- ³ Gran Sasso Science Institute, 67100 L'Aquila, Italy
- ⁴ INFN, Laboratori Nazionali del Gran Sasso, 67100 Assergi, AQ, Italy
- ⁵ Department of Physics and Engineering Physics Astronomy, Queen's University Kingston, Ontario, KIN K7L 3N6, Canada
- ⁶ Fermi National Accelerator Laboratory, Batavia, IL 60510, USA
- ⁷ Dipartimento di Fisica, Università di Milano-Bicocca, 20126 Milan, Italy
- ⁸ INFN, Sezione di Milano-Bicocca, 20126 Milan, Italy
- ⁹ Rigetti Computing, 2919 Seventh Street, Berkeley, CA 94710, USA
- ¹⁰ National Institute of Standards and Technology, Boulder, CO 80305, USA
- ¹¹ Department of Physics, University of Colorado, Boulder, CO 80309, USA
- ¹² Department of Electrical, Computer and Energy Engineering, University of Colorado Boulder, Boulder, CO 80309, USA
- ¹³ Dipartimento di Fisica, Sapienza Università di Roma, P.le Aldo Moro 2, 00185 Rome, Italy
- ¹⁴ Physics Department, University of Florida, Gainesville, USA
- ¹⁵ Center for Applied Physics and Superconducting Technologies, Northwestern University, Evanston, IL 60208, USA

Received: 29 November 2022 / Accepted: 4 January 2023 / Published online: 31 January 2023 © The Author(s) 2023

Abstract Radioactivity was recently discovered as a source of decoherence and correlated errors for the real-world implementation of superconducting quantum processors. In this work, we measure levels of radioactivity present in a typical laboratory environment (from muons, neutrons, and γ -rays emitted by naturally occurring radioactive isotopes) and in the most commonly used materials for the assembly and operation of state-of-the-art superconducting qubits. We present a GEANT-4 based simulation to predict the rate of impacts and the amount of energy released in a qubit chip from each of the mentioned sources. We finally propose mitigation strategies for the operation of next-generation qubits in a radio-pure environment.

Supplementary Information The online version contains supplementary material available at https://doi.org/10.1140/epjc/s10052-023-11199-2.

1 Introduction

The technology of macroscopic superconducting circuits offers many advantages in the development of quantum processors: ease in design, fabrication and operation, high fidelity, and fast gate-times. As proven by an increasing number of companies and research institutes, superconducting circuits are also primed to implement quantum entanglement, as they allow to inter-couple tens of qubits [1–4]. Both companies and research centers, are now aiming at a further scale-up in the number of entangled qubits. The work presented in this paper was done in the framework of the SQMS¹ (Superconducting Quantum Materials and System) Center, an international effort towards high-performance quantum computing.

Cosmic rays, as well as the decay of naturally radioactive isotopes present in the laboratory and in sample materials, can interact with the qubit chip, releasing energy. Energy deposits produce electron/holes charges and, subsequently, phonons. Both, charges and phonons, diffuse in the chip



^a e-mail: laura.cardani@roma1.infn.it

^b e-mail: francesco.dedominicis@gssi.it

^c e-mail: ambra.mariani@roma1.infn.it (corresponding author)

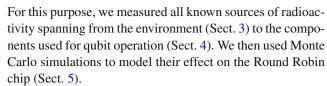
¹ https://sqms.fnal.gov.

with different footprints and interact with the qubits, resulting in a loss of coherence [5,6] and, if multiple qubits are involved, in correlated errors [6,7]. Phonons in particular can be absorbed by the superconductor, breaking Cooper pairs into dissipative quasiparticles. It is well known that quasiparticles can poison superconducting qubits [8–10] (including emerging cat-qubits [11]) and superconducting microwave resonators [12,13]. Moreover, new studies point to the existence of an interaction between ionizing radiation and the dominant qubit loss mechanism, namely two-level systems (TLSs) [14–21]. According to Thorbeck et al. [22], an energy deposit due to radioactivity causes frequency jumps in multiple TLSs, inducing fluctuations in the qubit lifetime and limiting the stability of the device.

Today, many groups are actively investigating strategies for the mitigation of the effects of radioactivity. Among the proposed ideas, we recall: (i) the suppression of radioactivity sources, benefiting from the experience of the scientific community developing low-radioactivity detectors for particle physics [23,24]; (ii) the development of "traps" surrounding the qubit to protect it from travelling phonons [25,26]; (iii) new processor design, in which matrices of qubits are deposited on chips decoupled from each other [27]; and (iv) assisting fault mitigation through the use of sensors located near physical qubits [28]. It is also worth mentioning emerging efforts in designing on-chip circulators to replace the present (bulky and full of ferrite) components and, at the same time, mitigate the tunneling of quasiparticles [29].

The work presented in this paper sets a first milestone for the SQMS Round Robin experiment, in which a standard multiqubit chip designed and fabricated by Rigetti Computing will be operated in multiple facilities belonging to the SQMS center (Boulder-Colorado, Fermilab and Northwestern University - Illinois, Rigetti - California, and the deep underground Gran Sasso Laboratories INFN-LNGS -Italy). These measurements entail long time, interleaved T1 and T2 scans in order to distinguish between decoherence channels and obtain reproducible performance metrics [30], with the final goal of disentangling all causes of decoherence and pinpointing new mitigation techniques. The contribution of INFN-LNGS, in particular, is establishing a lowradioactivity environment to test the SQMS prototypes, starting with the Round Robin one. In this framework, we foresee two main goals:

- Estimating the suppression of the radioactivity level in the "Round Robin" chip that can be obtained in the underground INFN-LNGS, in order to quantify the expected reduction before operating the prototype;
- Establishing a path to ensure that the radioactivity level of INFN-LNGS will satisfy the requirements for the future SQMS prototypes.



We highlight that, since radioactive interactions involve the qubit substrate (and not the qubit itself) [6,23], the obtained results apply to most qubits, with only minor adjustments due to the substrate volume. The effect of the energy deposits in the substrate on qubit performance, on the contrary, depends on the implementation of the superconducting circuit. The majority of superconducting qubits, indeed, is limited by TLS interactions. Radioactivity would become an issue only if this source of decoherence was mitigated. Nevertheless, some devices are already affected by the presence of quasiparticles or phonons [9,11,24,31], such as those produced by ionizing interactions in the substrate. As a consequence, suppressing radioactivity would directly enhance the performance of these qubits.

2 Setup of the prototype qubit

The Round Robin chip is a $325\,\mu\text{m}$ —thick, $11.9\times7\,\text{mm}^2$ high-resistivity silicon wafer hosting 16 transmon qubits (14 flux-tunable qubits, and two fixed-frequency qubits) with frequencies in the range 4.1– $4.9\,\text{GHz}$, and readout resonators with frequencies from 5.4 to $5.9\,\text{GHz}$. Since the vast majority of radioactive interactions involve the qubit substrate, we will only focus on the substrate and neglect the particular qubit design.

The chip is hosted in a copper box designed at NIST Boulder Laboratories, Colorado and its electrical connections are made using a printed circuit board (PCB). The box (Fig. 1) is mounted on a gold-plated copper holder, which is thermally anchored to the coldest stage of the dilution refrigerator (DR) at $\sim 10\,\mathrm{mK}$, and enclosed in a CryoPerm® magnetic shield (Fig. 2).

A variety of DRs will be used for these measurements at different institutes across the SQMS Center. Nevertheless, the materials used for the construction of the DUs and the dimensions of the vessels/components are rather similar, so the results we derived do not depend strongly on the particular design of the refrigerator. In the following simulations we modeled only the cryostat of INFN-LNGS. The full experimental setup including the cryostat and the corresponding simulated geometry is shown in Fig. 6.

Figure 3 shows a schematic view of the electronics components and readout to highlight their position compared to the qubit.



Eur. Phys. J. C (2023) 83:94 Page 3 of 10 94

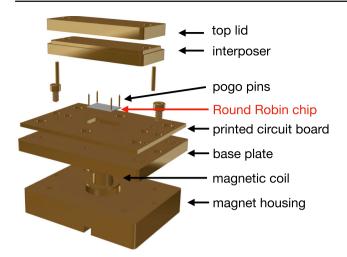


Fig. 1 Round Robin chip (gray) enclosed in its gold-plated copper box hosting the printed circuit board

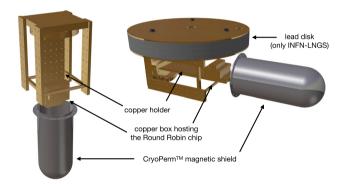


Fig. 2 Left: holder of the Round Robin chip. The copper box hosting the Round Robin chip, mounted on a gold-plated copper holder, is surrounded by a CryoPerm® magnetic shield. Gold-plated copper bars allow the installation of "cold" electronics components (not shown in the picture) in the proximity of the sample. The square copper element can be mounted on the bottom of the mixing chamber plate (the coldest point of the dilution refrigerator). Right: installation in INFN-LNGS, where the DU experimental space prevents the vertical configuration. The picture shows also the lead shield that can be added in the INFN-LNGS facility

Concerning the components that must be mounted close to the sample, we installed a Radiall² Cryogenic SP6T coaxial switch, a Low Noise Factory³ triple junction isolator and a Low Noise Factory single junction circulator. For each input line, we installed two cryogenic XMA⁴ attenuators (10 dB + 20 dB) and a 12 GHz low-pass filter provided by K&L Microwave.⁵ We anticipate that, in contrast to other facilities, the dilution refrigerator located in INFN-LNGS can be equipped with a lead shield at 10 mK, placed between the sample and the electronic components (Fig. 2).

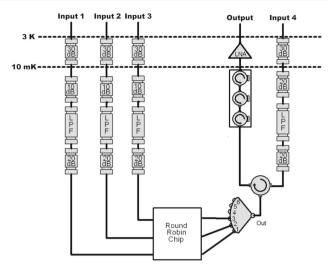


Fig. 3 Electronics setup for the Round Robin chip. The input lines are attenuated using 30 dB attenuators at 3 K and (10 + 20) dB attenuators at base temperature; moreover, they are filtered using 12 GHz low-pass filters ("LPF"). The output signal is amplified using a 4–8 GHz low noise cryogenic amplifier ("LNA"). A 6-channel cryogenic switch, low noise circulator and triple-junction isolator complete the readout

To connect these components to the room-temperature electronics, we used coaxial cables made of copper beryllium (from 300 K down to 3 K), and NbTi (from 3 K down to 10 mK). The final connections from the mixing chamber plate to the sample are done using copper amagnetic coaxial cables. The input lines are attenuated at 3 K using a 30 dB XMA attenuator. The output signal is amplified using a Low Noise Factory 4–8 GHz cryogenic amplifier.

3 Measurement of far radioactive sources

In the field of low-radioactivity detectors, it is common to distinguish between "far" sources of radioactivity and "close" sources of radioactivity, the main difference being that "close" sources cannot be shielded as they are in the proximity of the device or part of the device itself. As explained in this section, the INFN-LNGS offer a unique environment compared to other experimental sites involved in the Round Robin measurement, as the LNGS rock overburden naturally suppresses several "far" sources of radioactivity by orders of magnitude.

The most important sources of far radioactivity are cosmic rays, neutrons and γ -rays produced by radioactive decays in the laboratory environment. We underline that radioactive decays produce also other kind of particles, such as electrons and α particles. Nevertheless, the range of these particles in a medium-density material is of mm and μ m respectively, and thus they cannot penetrate the cryostat vessels or the copper box in which the qubit is hosted.



² https://www.radiall.com.

³ https://www.lownoisefactory.com.

⁴ https://www.xmacorp.com/.

⁵ https://www.klmicrowave.com.

94 Page 4 of 10 Eur. Phys. J. C (2023) 83:94

The main component of cosmic rays are muons. These ionizing particles arrive at sea level with an average energy of 4 GeV and a typical flux of about $1 \mu/\text{cm}^2/\text{min}$ [32]. The altitude of the experimental site, as well as the location of the laboratory within the site (e.g., if it is in the basement and thus shielded by the building itself) will impact the rate of muons reaching the device. The facilities involved in the Round Robin measurements are located about 0.05 km (Rigetti), 0.20 km (FNAL and Northwestern University) and 1.65 km (Boulder - CU) above sea level. Considering that the cryogenic facilities of CU, FNAL and Rigetti are located at the first floor of buildings that are not heavily shielded, we expect the muon flux in FNAL (CU) to be a factor 1.1 (1.3) larger than at Rigetti. The facility at Northwestern is being commissioned in the basement of the University, under five concrete floors. Thus, we expect a slightly lower muon flux compared to FNAL, despite the same elevation above sea level.

On the contrary, the cryogenic facility of INFN-LNGS is protected by a rock overburden of 1.4 km, enabling a suppression of the muon flux by six orders of magnitude [33,34].

The interaction of cosmic rays with our atmosphere also produces neutrons. Their flux at the surface extends from thermal energies (meV) to GeV, with an intensity that varies with altitude, geomagnetic field, and solar magnetic activity ([35] and references therein). In shallow and underground sites, low energy neutrons (below 10 MeV) are produced both by spontaneous fission and (α,n) reactions, while fast neutrons are produced by nuclear reactions induced by residual cosmic ray muons in the rock or in the experimental apparatus.

We measured the flux of environmental neutrons in the energy region (0–20) MeV with a DIAMON neutron spectrometer [36], a portable detection system that provides real-time neutron spectrometry. The spectrum obtained *above ground* at INFN-LNGS is shown in Fig. 4. We measured a flux of 0.018 n/cm²/s, with an average experimental data uncertainty of 7%.

The same neutron measurement was performed in other three laboratories located in different cities, obtaining negligible variations in the spectral shape, and a total flux spanning from 0.010 to 0.018 n/cm²/s. Thus, for the following study we will use the energy spectrum shown in Fig. 4 with an average flux of (0.014 ± 0.004) n/cm²/s, for all the sites located above ground.

On the contrary, the neutron flux in the underground INFN-LNGS has been measured to be several orders of mag-

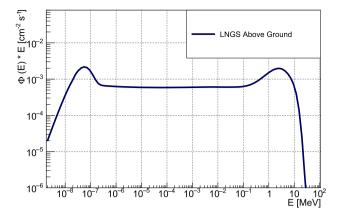


Fig. 4 Spectrum of neutrons measured above-ground at the INFN-LNGS. The total flux is $0.018\,\text{n/cm}^2/\text{s}$, with an average experimental data uncertainty of 7%

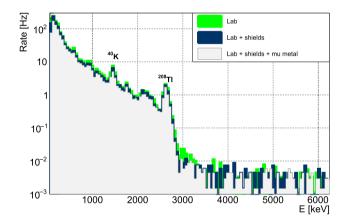


Fig. 5 Energy spectrum of the γ -ray flux measured using a 3" portable NaI spectrometer. Green: original spectrum collected in a laboratory of cryogenic detectors (Italy). Blue: same spectrum inside the cryostat thermal and vacuum vessels. Gray: same spectrum as (blue) but adding also the mu-metal magnetic shield

nitude lower than the atmospheric neutron flux [37], on the order of 10^{-6} n/cm²/s, with the rate of muon-induced neutrons from two to three orders of magnitude lower than the rate of neutrons produced by nuclear reactions in the mountain rocks.

Environmental γ -rays have average fluxes of few γ /cm²/s and typical energies lower than 2.6 MeV (one of the γ -rays produced by the decay of the 208 Tl isotope). Figure 5 shows the γ -rays spectrum measured with a 3" portable NaI spectrometer in a laboratory of cryogenic detectors (Roma, Italy).

As in the case of neutrons, the shape of the spectrum is expected to be more or less the same in the different Round Robin locations. Indeed, most of the radioactive isotopes naturally present in the environment belong to the ²³²Th- or ²³⁸U-chains (with the exception of the 1.4 MeV peak due to the primordial radionuclide ⁴⁰K). On the other hand, the rate of each peak depends on the specific contamination of the laboratory, and on the total shielding. For this work, we



⁶ In a heavily shielded laboratory located at the Chooz Nuclear Power Station (France), we measured a neutron flux of 0.005 n/cm²/s. In a laboratory located in the Roma University – Sapienza basement, we measured 0.007 n/cm²/s. However, none of the laboratories involved in the Round Robin measurements is shielded by the building itself so we did not use these results.

Eur. Phys. J. C (2023) 83:94 Page 5 of 10 94

measured the absolute flux in the experimental hall of INFN-LNGS and in other above ground laboratories, obtaining (1.0 \pm 0.5) γ /cm²/s and (2.5 \pm 0.5) γ /cm²/s respectively.⁷

Finally, Fig. 5 shows three measurements performed in the same site to evaluate the effect of the DR itself on the chip. Typical DRs, indeed, include one or two vacuum cans, and thermal vessels to protect the samples from thermal radiation. In addition, they can be equipped with magnetic shields. We performed the same measurement in three different scenarios: (i) without vessels, (ii) closing the DR with its thermal vessels and vacuum cans, and (iii) adding also a mu-metal magnetic shield. The γ -ray flux, obtained by integrating the gamma rate over the entire energy range, is reduced by a factor of 25% and 33% inside the vessels and inside the vessels + mu-metal respectively. Our simulation (Sect. 5) accounts for such effects.

4 Measurement of close radioactive sources

Close sources of radioactivity comprise cables, electronic components, and other materials that cannot be placed far away from the chip. In this work, we focus on the components that will be used for the Round Robin measurements. Nevertheless, the vast majority of these components are common in many experimental facilities and, thus, this information can be used to predict background contributions in different qubit prototypes. Our list includes:

- A printed circuit board (PCB) weight: ~ 7 g. The PCBs used for the Round Robin were produced by San Francisco Circuits. They consisted of a 1.575×2.205 cm², 0.157 cm thick 3-layer FR408HR (a high performance laminate for multi-layer Printed Wiring Board applications, where maximum thermal performance is required); We also measured a new type of non-magnetic PCB, that however was not mounted in the current prototype of the Round Robin chip and is indicated with A* in Table 1;
- B gold-plated copper box (total mass: 0.4 kg) and a gold-plated copper holder (2.2 kg) used to thermally anchor the qubit to the mixing chamber plate as shown in Fig. 2;
- C "cold" magnetic shield, consisting in a hollow CryoPerm® cylindrical shield with 78.8 mm diameter, 193.5 mm height and 1.0 mm thickness weight: \sim 475 g;
- D Intelliconnect non-magnetic SMA adapters weight: \sim 10 g;
- E Intelliconnect non-magnetic copper coaxial cables (diameter: 2.19 mm, length: 25 cm);
- F Radiall SP6T cryogenic switch weight: $\sim 165 \, \mathrm{g}$;

- G 4–12 GHz Low Noise Factory circulator (mod. CIC4_12A), not used in the Round Robin measurements weight: ~50 g;
- H 4–12 GHz Low Noise Factory dual-junction circulator (mod. CICIC4_12A), not used in the Round Robin measurements weight: ~ 97 g;
- I 4–8 GHz Low Noise Factory triple-junction isolator (mod. ISISISC4 8A) weight: ~ 145 g;
- J 10–30 dB XMA attenuators (mod. 2082-6418-10-CRYO, 2082-6418-20-CRYO, and 2082-6418-30-CRYO) weight: \sim 5 g each;
- K K&L Low Pass Filters (mod. 3L250-12240/T20000 and 3L250-8160/T20000) weight: $\sim 15 \,\mathrm{g}$ each;
- L COAX⁸ NbTi superconducting coaxial cables running from 4 to 15 mK (mod SC-219/50-NbTi-NbTi).

Other components are placed at the 4K stage and thus, their contribution to the qubit counting rate is mitigated by their distance from the sample:

- M 4–8 GHz Low Noise Factory cryogenic amplifier weight: ~ 17 g;
- N COAX copper-beryllium coaxial cables running from $300\,\mathrm{K}$ to $4\,\mathrm{K}$ (mod SC-119/50-B-B).

Finally, we report the results for a thermally conductive epoxy glue (Stycast®, O*) and DOW CORNING cryogenic grease (P*). These materials are widely used for qubit application and their radioactive content is of general interest. However, we did not make use of them for the Round Robin chip.

We investigated the radiopurity of these components using the γ -spectrometric techniques detailed in the Supplemental Materials. The results of the screening, reported in Table 1, show that the PCBs (A) contain a significant amount of natural radioactivity. This was expected, due to the PCBs composition (i.e. glass fiber) [38–40]. We also include a possible alternative type of a-magnetic PCB (A*), consisting of copper (three layers, for a total thickness of 87 μ m) interleaved by dielectric layers (hydrocarbon ceramic laminates) produced by ROGERS Corporation⁹ under the codes 4350B (two layers, for a total thickness of 420 μ m) and 4450F (a single layer, 95 μ m thick). As shown in Table 1, these PCBs (A*) were about a factor of 3 cleaner, from the radioactivity point of view, compared to those selected for the Round Robin chip.

On the contrary, the materials dominant in weight and closest to the chip – copper (B) and CryoPerm (C) – feature a good radiopurity level. Nevertheless, these components are



⁷ This value is representative for a "typical" laboratory environment. In buildings made of particulars materials (such as tuff rock) the γ -ray flux can be a factor of three higher.

⁸ http://www.coax.co.jp.

https://www.rogerscorp.com.

94 Page 6 of 10 Eur. Phys. J. C (2023) 83:94

Table 1 Bulk contamination – values and 90% C.L. upper limits – of the materials located in the proximity of the qubit at the 15 mK stage (A–L) and 4 K (M,N) stage. Items indicated with (*) were measured for general interest but will not be used in the Round Robin measurements. The full results are reported in the Supplemental Material and summa-

rized in this table: for the ²³²Th decay chain, we quote the maximum activity between the ones measured for ²²⁸Ra and ²²⁸Th; concerning ²³⁸U and daughters, we quote the ²²⁶Ra activity, being representative of the most worrisome part of the decay chain

Component	²³² Th (mBq/kg)	²³⁸ U (mBq/kg)	²³⁵ U (mBq/kg)	40 K (mBq/kg)	¹³⁷ Cs (mBq/kg)
\overline{A}	$(18,000 \pm 1000)$	$(11,500 \pm 400)$	(710 ± 110)	$(12,000 \pm 1000)$	< 30
A*	(5410 ± 330)	(4200 ± 200)	(230 ± 50)	(4200 ± 500)	< 40
B	< 1.5	< 1.2	< 4	< 9	< 0.6
C	< 8.4	< 8.3	< 8.4	< 35	< 2.7
D	(46 ± 13)	(42 ± 10)	(70 ± 30)	(240 ± 90)	< 10
E	(54 ± 12)	(44 ± 11)	(34 ± 17)	(740 ± 130)	< 12
F	(1880 ± 100)	(1340 ± 60)	(130 ± 30)	(2200 ± 300)	< 11.2
G	< 310	< 330	< 410	< 2000	< 60
H^*	< 250	< 380	< 380	< 2600	< 60
I^*	< 190	< 240	< 220	< 2000	< 50
J	< 52	(200 ± 20)	< 47	< 140	< 13
K	(23 ± 4)	< 9.1	(60 ± 10)	< 100	< 1.9
L	< 750	< 1000	< 380	< 7000	< 230
M	< 890	< 1000	< 850	< 10000	< 210
N	(240 ± 40)	< 78	(350 ± 90)	< 500	< 20
O^*	(53 ± 4)	(9400 ± 900)	(350 ± 30)	(290 ± 40)	< 2.2
P^*	< 10	< 11	< 4.5	< 87	< 5

Table 2 Activity concentration – values and 90% C.L. upper limits – for short- and long-lived radioisotopes produced by cosmogenic activation in the materials of the detector setup

Component	⁶⁰ Co (mBq/kg)	⁵⁸ Co (mBq/kg)	⁵⁷ Co (mBq/kg)	⁵⁴ Mn (mBq/kg)
В	(0.6 ± 0.3)	(4 ± 1)	_	(2.4 ± 0.8)
C	< 3.7	(14 ± 5)	(20 ± 7)	_
D	(51 ± 8)	_	_	_
K	(5 ± 1)	_	_	_

the ones most subject to cosmogenic activation. Indeed, this analysis confirms the presence of cosmogenic radioisotopes – ⁵⁴Mn, ⁵⁷Co, ⁵⁸Co, and ⁶⁰Co – produced above-ground by cosmic ray interaction with the materials [41,42]. The results are reported in Table 2.

5 Effects on the qubit chip

The geometry of the setup described in Figs. 1 and 2 was implemented in our GEANT-4 based simulation framework [43]. We also implemented the copper holder, the magnetic shield, and a simplified version of the dilution refrigerator, including its internal and external lead shields (Fig. 6). The Round Robin chip was considered the only active volume, where we stored tracks and recorded the energy deposited by each simulated interaction. In the Monte Carlo simulations two different approaches were chosen for far and close sources.

For the far sources (γ -rays, muons and neutrons), we generated primary particles within the laboratory environment. More in detail, for neutrons and γ -rays we generated the events uniformly distributed on the surface of a cylinder enclosing the cryostat, according to the measured energy spectra (Figs. 4, 5) and with isotropic momentum distribution. Concerning muons, in order to reproduce the actual angular distribution, we randomly generate positions on an hemisphere around the cryostat according to a \cos^2 distribution. For every sampled position, we generate muons perpendicularly from a $(120 \times 120) \, \mathrm{cm}^2$ tangent-plane to the hemisphere.

We then estimate the rate of interactions in the chip for each of these sources by scaling the number of recorded events to the flux measurements (Sect. 3 and references



¹⁰ In INFN-LNGS the angular distribution is affected by the rock overburden and cannot be accurately described by the cos² distribution. This does not impact significantly the result.

Eur. Phys. J. C (2023) 83:94 Page 7 of 10 94

Fig. 6 Experimental setup as implemented in the simulation. The chip inside the copper box is the only active volume. The copper box, holder and magnetic shield were imported from the CAD file, while for the dilution refrigerator we implemented a simplified version

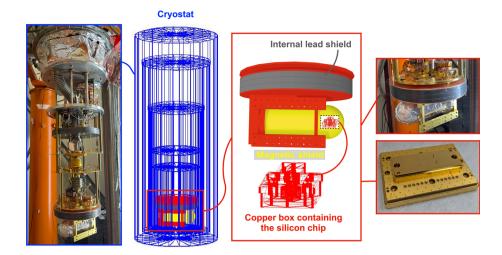


Table 3 Interaction rate in the substrate of the Round Robin chip. For the above-ground facilities we assumed a "standard" γ -ray flux of (2.5 \pm 0.5) γ /cm²/s and a muon flux of 1 μ /cm²/min. For the LNGS facility

we used the measured flux of $(1.0 \pm 0.5) \gamma/\text{cm}^2/\text{s}$ and we evaluated the suppression factor due to the presence of lead shields

Source	"Standard" (mHz)	LNGS (mHz)	LNGS Ext. Shield (mHz)	LNGS Full Shield (mHz)
Lab γ-rays	(18 ± 4)	(7.0 ± 3.5)	(0.10 ± 0.05)	(0.07 ± 0.04)
Muons	(10.0 ± 0.6)	$<10^{-5}$	<10 ⁻⁵	<10 ⁻⁵
Neutrons	(0.15 ± 0.05)	$< 10^{-4}$	$< 10^{-4}$	$< 10^{-4}$

therein), both for the above-ground laboratories ("standard") and the underground INFN-LNGS.

Simulations were done considering the three possible INFN-LNGS shielding setups: no shield at all, only the external shield surrounding the cryostat (10 cm thick wall of lead bricks), and the so-called "full" shield configuration, with both the external lead shield and the inner lead shield (3 cm thick lead disk between the chip and the mixing chamber). The results are summarized in Table 3.

Concerning above-ground facilities, the highest interaction rate in the Round Robin chip is produced by γ -rays from laboratory radioactivity, resulting in a rate of (18 \pm 4) mHz. The rate in INFN-LNGS is only (7.0 ± 3.5) mHz, due to intrinsically lower content of natural radioactivity in the Gran Sasso rock. This rate is further suppressed through the combination of the external and internal lead shields. More in detail, the external shield alone and the internal shield alone abate the γ -ray flux by a factor 70 and 2 respectively. Their combination offers a flux suppression by a factor 100, for a final value of (0.07 ± 0.04) mHz. This suppression factor is based on the assumption that the external lead shield is an ideal cylinder 80 cm tall, 10 cm thick. Less tight geometries could limit the suppression capability of the external shield. The intrinsic radioactivity of the lead shields could in principle generate interactions in the Round Robin chip. To check this, we simulated a typical contamination of ²¹⁰Pb with an activity concentration of 100 Bq/kg in the internal lead shield (the most worrisome one, due to its proximity with the sample) and obtained a negligible interaction rate in the Round Robin chip of 0.01 mHz.

In conclusion, this shielding system is very effective in suppressing the dominant radioactive source and could be easily installed also in above ground facilities.¹¹

The second most significant contribution to the rate of the Round Robin chip ((10.0 \pm 0.6) mHz in the Rigetti laboratories) is due to muons. Despite their lower rate, muons are considered more worrisome than γ -rays, as they produce long tracks potentially affecting more qubits lying on the same chip [6].

The only viable strategy to suppress muons is moving the facility to an underground site. As shown in Table 3, INFN-LNGS offers an extremely high reduction of cosmic rays; nevertheless, effective suppression factors can be obtained even in shallow sites. As an instance, the shallow underground laboratory NEXUS (Northwestern Experimental Underground Site at Fermilab, Illinois, US) is shielded by only 100 m of soil/gravel, and yet offers a suppression by three orders of magnitude of atmospheric muons [44,45].

An alternative strategy may consist of tagging muons (instead of suppressing them) in order to identify and reject operations done while these particles are traversing the chip.

¹¹ The shielding capability in above-ground facilities would be slightly less effective due to the presence of high energy γ -rays produced by cosmic rays.



Table 4 Interaction rate in the substrate of the Round Robin chip. The copper box and holder were split in this Table to highlight that, due to its proximity to the qubit, the contribution of the box is more important than the one of the holder, despite the same level of radiopurity. Some components are highlighted in bold to remind that usually these components are placed in large quantity in the dilution refrigerator (and the rate should be scaled accordingly)

Component	Description	Rate (mHz)
\overline{A}	PCB	4.52 ± 0.04
B	Box	$[1-6] \times 10^{-3}$
B^*	Holder	$[2-4] \times 10^{-4}$
C	Magnetic shield	$[2-9] \times 10^{-4}$
D	SMA	$(2 \pm 0.4) \times 10^{-5}$
E	Cu coax cables	$(3 \pm 0.6) \times 10^{-5}$
F	Cryogenic switch	$(1.0 \pm 0.2) \times 10^{-2}$
G	Circulator	$< 8 \times 10^{-4}$
H^*	Dual-junct. circulator	$< 2 \times 10^{-3}$
I^*	Triple-junct. isolator	$< 2 \times 10^{-3}$
J	Attenuators	$[0.5-1] \times 10^{-5}$
K	Low pass filters	$(1 \pm 0.2) \times 10^{-5}$
L	NbTi cables	$< 4 \times 10^{-4}$
M	Cryogenic amplifier	$< 2 \times 10^{-5}$
N	Cu-Be cables	1×10^{-6}

The drawback of this approach is that a compact and efficient external muon veto to be operated above-ground would feature a trigger rate as high as hundreds of Hz [46]. More sophisticated vetoes at cryogenic temperature could guarantee a similar efficiency while diminishing the trigger rate (and thus, the associated dead time).

We also recall that other groups are proposing to suppress the devastating effects of muons at the chip level by proposing novel distributed error correction schemes [47].

Finally, we observe that neutrons, that are the main concern for classical computers [48], have a negligible impact rate in typical superconducting qubits.

"Close" sources of radioactivity were simulated by generating the radioactive decays of the relevant isotopes uniformly distributed within the volume of each component (A-N). The location of every component in the simulated geometry was assumed as the typical place where the given item is usually mounted in the cryostat. Due to the proximity of some components to the Round Robin chip (PCB and copper box in particular), we simulated also β - and α -particles, in addition to γ -rays. We then estimated the rate of interactions in the chip from each component by scaling the number of recorded events to the measured activity concentrations reported in Table 5 in the Supplemental Material (and summarized in Table 1).

The results, summarized in Table 4, show that the vast majority of electronics components are not worrisome for

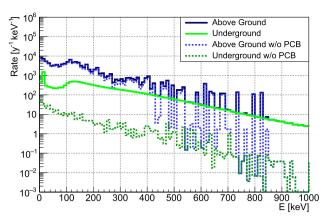


Fig. 7 The rate of interactions in the Round Robin chip is reported as a function of the energy that they deposit in the silicon chip. The majority of interactions deposit tens/hundreds of keV. Solid lines: total rate expected in a typical above ground laboratory (blue) and in the INFN-LNGS underground site (green). Dashed lines show what would be the rate in the two cases without the PCB contribution. Above ground, the effect of the PCB is almost negligible, as the rate is dominated by far radioactive sources. On the contrary, in the deep underground INFN-LNGS the PCB is the major contribution to the interaction rate. As a consequence, substituting the PCB with a more radio-pure one would allow to abate the rate of interactions in the chip

qubit measurements, with the exception of the PCB. Being very close to the qubit and, at the same time, not very radiopure, the PCB constitutes an irreducible source of interactions in the Round Robin chip. This problem is well known to particle physicists developing detectors for rare events searches, as both fiberglass and ceramic (typical materials used for the PCB multilayers) contain usually levels of radioactivity comparable to those in rocks and soils. The presence of the PCB is the ultimate limit to the improvement of the radioactivity level in the INFN-LNGS laboratory environment, compared to above-ground sites.

In Fig. 7, we report the rate of events as a function of the energy that they deposit in the Round Robin chip. In this plot we distinguish the results expected for the above-ground laboratories and for INFN-LNGS, where the overall rate is expected to be about an order of magnitude lower. In both cases we show the results with and without the PCB (the only relevant contribution of the "close" sources), to show the improvement that might be obtained by substituting this component with a more radio-pure one.

We want to emphasize that our screening measurements (Sect. 4) are not sensitive to fast decaying isotopes, i.e. radioisotopes that are activated above-ground and decay within seconds or minutes. In principle, these radioisotopes could further increase the interaction rate in the chip when the Round Robin is operated above-ground. Nevertheless, we simulate the effect of the cosmogenic activation of copper (the most "massive" material in the setup) using the ACTIVIA software [49] and we obtained a negligible rate



Eur. Phys. J. C (2023) 83:94 Page 9 of 10 94

in the chip of 0.05 mHz. Furthermore, some of the components (G, H, I, M) were measured also above ground, finding no evidence of activated isotopes within the reported sensitivities.

This study proves that the radioactivity suppression obtained in INFN-LNGS can be considered satisfactory for the Round Robin measurements, as it allows us to search for macroscopic differences in the prototype performance.

For a long-term goal of reaching qubits lifetime of a seconds (the goal of the 3-D SQMS architecture), it is likely that further improvements will be required. Indeed, with the measured contamination levels in INFN-LNGS, the probability of observing a radioactivity-induced event in a secondlong time window, amounts to 0.4%. Nevertheless, such a rate was estimated using the geometry of the Round Robin chip. A chip hosting 256 qubits will likely demand a larger substrate and thus be prone to a higher rate of interactions, requiring lower radioactivity levels. The PCB in particular could become the main issue in suppressing the interaction rate. As explained in Sect. 4, we already measured the content of natural radioactivity of a new type of non-magnetic PCB, obtaining a factor \sim 3 improvement compared to the Round Robin PCBs (Table 1). A further improvement, if needed, will require a dedicated R&D activity.

6 Conclusions

In this work, we measured the radioactive content of all the components that are commonly used in the characterization of superconducting qubits. We considered a particular case (the Round Robin chip of the SQMS center) to run Monte Carlo simulations predicting the impact of each radioactive source on the qubit chip.

We conclude that the overall interaction rate is dominated by γ -rays radioactivity of the laboratory environment, that can be suppressed using lead shields.

The second contribution comes from muons. Even if muon-vetoes, or on-chip mitigation strategies can be envisioned, the most effective abatement strategy for these interactions consists in operating the prototypes in underground sites, such as the INFN-LNGS for the SQMS center.

Finally, the setup components resulted generally radiopure, with the exception of PCBs, that demand a dedicated optimization. Our studies already identified an intermediate solution, allowing to suppress the interaction rate coming from the PCBs by roughly a factor of 3, but other solutions may be necessary for next-generation quantum processors.

Acknowledgements The authors thank the Director and technical staff of the Laboratori Nazionali del Gran Sasso. We are grateful to A. Girardi for the controller of the cryogenic switch, M. Guetti for the help with the whole cryogenic facility (including its continuous upgrades), and M. Iannone for the design and construction of the chip holder and read-

out. We are also grateful to the LNGS Computing and Network Service for computing resources and support on U-LITE cluster at LNGS. We thank F. Cappella and G. Pessina for the stimulating discussions on the simulator of muons and on the radioactivity of the PCBs, respectively. The work was supported by the U.S. Department of Energy, Office of Science, National Quantum Information Science Research Centers, Superconducting Quantum Materials and Systems Center (SQMS) under the contract no. DE-AC02-07CH11359, and by the Italian Ministry of Research under the PRIN Contract no. 2020h51338 ("Thin films and radioactivity mitigation to enhance superconducting quantum processors and low temperature particle detectors"). Finally, we are grateful to Raffele Tripiccione (Lele) for his tireless effort in building this collaboration.

Data Availability Statement This manuscript has no associated data or the data will not be deposited. [Authors' comment: Data will be provided upon reasonable request.]

Open Access This article is licensed under a Creative Commons Attribution 4.0 International License, which permits use, sharing, adaptation, distribution and reproduction in any medium or format, as long as you give appropriate credit to the original author(s) and the source, provide a link to the Creative Commons licence, and indicate if changes were made. The images or other third party material in this article are included in the article's Creative Commons licence, unless indicated otherwise in a credit line to the material. If material is not included in the article's Creative Commons licence and your intended use is not permitted by statutory regulation or exceeds the permitted use, you will need to obtain permission directly from the copyright holder. To view a copy of this licence, visit http://creativecommons.org/licenses/by/4.0/.

Funded by SCOAP³. SCOAP³ supports the goals of the International Year of Basic Sciences for Sustainable Development.

References

- Rigetti aspen-m-2 features. https://www.rigetti.com/ what-we-build. Accessed 2 Dec 2022
- F. Arute et al., Nature 574(7779), 505 (2019). https://doi.org/10. 1038/s41586-019-1666-5
- Y. Wu et al., Phys. Rev. Lett. (2021). https://doi.org/10.1103/ physrevlett.127.180501
- J. Chow et al., Ibm quantum breaks the 100-qubit processor barrier (2021). https://research.ibm.com/blog/ 127-qubit-quantum-processor-eagle. Accessed 16 Nov 2021
- A.P. Vepsäläinen et al., Nature 584(7822), 551 (2020). https://doi. org/10.1038/s41586-020-2619-8
- C.D. Wilen et al., Nature 594(7863), 369 (2021). https://doi.org/ 10.1038/s41586-021-03557-5
- M. McEwen et al., Nat. Phys. 18(1), 107 (2021). https://doi.org/ 10.1038/s41567-021-01432-8
- G. Catelani et al., Phys. Rev. B (2011). https://doi.org/10.1103/ physrevb.84.064517
- I.M. Pop et al., Nature 508(7496), 369 (2014). https://doi.org/10. 1038/nature13017
- U. Vool et al., Phys. Rev. Lett. (2014). https://doi.org/10.1103/ physreylett.113.247001
- C. Berdou et al., One hundred second bit-flip time in a twophoton dissipative oscillator (2022). https://doi.org/10.48550/ ARXIV.2204.09128. arXiv:2204.09128
- U. Patel et al., Phys. Rev. B (2017). https://doi.org/10.1103/ physrevb.96.220501
- S.A.H. de Rooij et al., Phys. Rev. B 104, L180506 (2021). https://doi.org/10.1103/PhysRevB.104.L180506



94 Page 10 of 10 Eur. Phys. J. C (2023) 83:94

- P.V. Klimov et al., Phys. Rev. Lett. 121, 090502 (2018). https://doi. org/10.1103/PhysRevLett.121.090502
- S.E. de Graaf et al., Phys. Rev. Lett. 118, 057703 (2017). https://doi.org/10.1103/PhysRevLett.118.057703
- J. Burnett et al., Nat. Commun. 5(1), 4119 (2014). https://doi.org/ 10.1038/ncomms5119
- A. Bilmes et al., Sci. Rep. 10(1), 3090 (2020). https://doi.org/10. 1038/s41598-020-59749-y
- J. Lisenfeld et al., npj Quantum Inf. 5(1), 105 (2019). https://doi. org/10.1038/s41534-019-0224-1
- J.M. Martinis et al., Phys. Rev. Lett. 95, 210503 (2005). https://doi.org/10.1103/PhysRevLett.95.210503
- C. Müller et al., Rep. Prog. Phys. 82(12), 124501 (2019). https://doi.org/10.1088/1361-6633/ab3a7e
- C.R.H. McRae et al., Rev. Sci. Instrum. 91(9), 091101 (2020). https://doi.org/10.1063/5.0017378
- T. Thorbeck et al., Tls dynamics in a superconducting qubit due to background ionizing radiation (2022). https://doi.org/10.48550/ ARXIV.2210.04780. arXiv:2210.04780
- L. Cardani et al., Nat. Commun. (2021). https://doi.org/10.1038/ s41467-021-23032-z
- D. Gusenkova et al., Appl. Phys. Lett. 120(5), 054001 (2022). https://doi.org/10.1063/5.0075909
- F. Henriques et al., Appl. Phys. Lett. 115(21), 212601 (2019). https://doi.org/10.1063/1.5124967
- J.M. Martinis, npj Quantum Inf. 7(1), 90 (2021). https://doi.org/ 10.1038/s41534-021-00431-0
- A. Gold et al., npj Quantum Inf. 7(1), 142 (2021). https://doi.org/ 10.1038/s41534-021-00484-1
- J.L. Orrell, B. Loer, Phys. Rev. Appl. 16, 024025 (2021). https://doi.org/10.1103/PhysRevApplied.16.024025
- D.T. Le et al., Phys. Rev. Res. 3, 043211 (2021). https://doi.org/ 10.1103/PhysRevResearch.3.043211
- C.R.H. McRae, G.M. Stiehl, H. Wang, S.X. Lin, S.A. Caldwell, D.P. Pappas, J. Mutus, J. Combes, Appl. Phys. Lett. 119(10), 100501 (2021). https://doi.org/10.1063/5.0060370
- L. Grünhaupt, M. Spiecker, D. Gusenkova, N. Maleeva, S.T. Skacel, I. Takmakov, F. Valenti, P. Winkel, H. Rotzinger, W. Wernsdorfer, A.V. Ustinov, I.M. Pop, Nat. Mater. 18(8), 816 (2019). https://doi. org/10.1038/s41563-019-0350-3
- M. DePascale et al., J. Geophys. Res. 98(A3), 3501 (1993). https://doi.org/10.1029/92JA02672
- M. Ambrosio et al., Phys. Rev. D 52, 3793 (1995). https://doi.org/ 10.1103/PhysRevD.52.3793
- M. Aglietta et al., Phys. Rev. D 58, 092005 (1998). https://doi.org/ 10.1103/PhysRevD.58.092005
- M. Gordon et al., Nucl. Sci. IEEE Trans. 51, 3427 (2005). https://doi.org/10.1109/TNS.2004.839134. https://ieeexplore.ieee.org/document/1369506
- A. Pola et al., Nucl. Instrum. Methods. Phys. Res. A 969, 164078 (2020). https://doi.org/10.1016/j.nima.2020.164078

- A. Best et al., Nucl. Instrum. Methods. Phys. Res. A 812, 1 (2016). https://doi.org/10.1016/j.nima.2015.12.034
- 38. E. Aprile et al., Eur. Phys. J. C 77(12), 890 (2017). https://doi.org/ 10.1140/epjc/s10052-017-5329-0
- E. Armengaud et al., J. Instrum. 12(08), P08010 (2017). https://doi.org/10.1088/1748-0221/12/08/P08010
- H. Ito et al., Nucl. Instrum. Methods Phys. Res. Sect. A Accelerators Spectrom. Assoc. Equip. 953, 163050 (2020). https://doi.org/10.1016/j.nima.2019.163050
- 41. L. Baudis et al., Eur. Phys. J. C **75**(10), 485 (2015). https://doi.org/10.1140/epjc/s10052-015-3711-3
- S. Cebrián, Int. J. Mod. Phys. A 32(30), 1743006 (2017). https://doi.org/10.1142/S0217751X17430060
- 43. S. Agostinelli et al., Nucl. Instrum. Methods. Phys. Res. A **506**(3), 250 (2003). https://doi.org/10.1016/S0168-9002(03)01368-8
- R. Ren et al., Phys. Rev. D (2021). https://doi.org/10.1103/ physrevd.104.032010
- M. Battaglieri et al., Us cosmic visions: new ideas in dark matter 2017: community report (2017). https://doi.org/10.48550/ARXIV. 1707.04591. arXiv:1707.04591
- V. Wagner et al., J. Instrum. 17(05), T05020 (2022). https://doi. org/10.1088/1748-0221/17/05/t05020
- Q. Xu et al., Distributed quantum error correction for chip-level catastrophic errors (2022). https://doi.org/10.48550/ARXIV.2203. 16488. arXiv:2203.16488
- R. Baumann, IEEE Des. Test Comput. 22(3), 258 (2005). https://doi.org/10.1109/MDT.2005.69. https://ieeexplore.ieee.org/document/1438282/references#references
- J. Back, Y. Ramachers, Nucl. Instrum. Methods. Phys. Res. A 586(2), 286 (2008). https://doi.org/10.1016/j.nima.2007.12.008
- A. Borio di Tigliole, A. Cammi, D. Chiesa, M. Clemenza, S. Manera, M. Nastasi, L. Pattavina, R. Ponciroli, S. Pozzi, M. Prata, E. Previtali, A. Salvini, M. Sisti, Progr. Nucl. Energy 70, 249 (2014). https://doi.org/10.1016/j.pnucene.2013.10.001
- M. Laubenstein, I. Lawson, Front. Phys. 8, 506 (2020). https://doi. org/10.3389/fphy.2020.577734
- F. Bellini et al., Eur. Phys. J. C 81(2), 106 (2021). https://doi.org/10.1140/epjc/s10052-021-08918-y. https://epjc.epj.org/articles/epjc/abs/2021/02/10052_2021_Article_8918/10052_2021_Article_8918.html
- M. Beretta, M. Biassoni, L. Gironi, M. Maino, M. Nastasi, L. Pagnanini, S. Pozzi, Eur. Phys. J. Plus 136(1), 89 (2021). https://doi.org/10.1140/epjp/s13360-020-01068-2
- F.K.M. Heisel, H. Simgen, Appl. Radiat. Isot. 67(5), 741 (2009). https://doi.org/10.1016/j.apradiso.2009.01.028. https://pubmed.ncbi.nlm.nih.gov/19251428/

